

Decay behavior and optical parameter identification for spatial-frequency domain imaging by the radiative transport equation

MANABU MACHIDA,^{1,*} ^(D) YOKO HOSHI,¹ KEIICHIRO KAGAWA,² AND KAZUKI TAKADA³

¹Institute for Medical Photonics Research, Hamamatsu University School of Medicine, Hamamatsu 431-3192, Japan ²Research Institute of Electronics, Shizuoka University, Hamamatu 432-8011, Japan ³Graduate School of Integrated Science and Technology, Shizuoka University, Hamamatu 432-8011, Japan *Corresponding author: machida@hama-med.ac.jp

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The decay behavior of specific intensity is studied for spatial-frequency domain imaging (SFDI). It is shown using the radiative transport equation that the decay is given by a superposition of different decay modes, and the decay rates of these modes are determined by spatial frequencies and Case's eigenvalues. This explains why SFDI can focus on shallow regions. The fact that light with nonzero spatial frequency rapidly decays makes it possible to exclusively extract optical properties of the top layer of a layered medium. We determine optical properties of the top layer of a solid phantom. This measurement is verified with different layered media of numerical phantoms. © 2020 Optical Society of America under the terms of the OSA Open Access Publishing Agreement

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1. INTRODUCTION

In near-infrared spectroscopy, light illumination in the spatialfrequency domain has been developed as a tool that is as concise as continuous-wave illumination and as informative as frequency-domain illumination [1]. Spatial-frequency domain imaging (SFDI) is capable of determining both absorption and scattering coefficients from time-independent measurements. SFDI is used mainly to extract optical properties at depths on the order of millimeters. SFDI was used for imaging skin flap oxygenation during reconstructive breast surgery [2]. It was also used to record biochemical compositional changes in a port wine stain after laser therapy [3]. Tissue optical properties of a human volar forearm were estimated by SFDI [4]. Burn wounds were examined by SFDI coupled with laser speckle imaging [5]. See a recent review by Angelo *et al.* [6] and references therein.

The abovementioned works show that SFDI can exclusively study shallow regions near the skin. In SFDI, spatially modulated incident beams rapidly decay in biological tissue [7]. In this paper, we further investigate this feature of SFDI. This property of SFDI is advantageous when we are interested in measuring optical properties of shallow regions. Even when the thickness is thin, we can assume the half space, which is unbounded in the depth direction. In this paper, we identify optical properties of the top layer of a layered medium. It is not possible to extract optical properties of the top layer in the standard setting of nearinfrared spectroscopy, in which optical fibers are attached on the top of a layered medium, because near-infrared light propagates not only in the top layer but reaches deeper layers.

In this paper, the ability to extract optical properties of the top layer is tested by different numerical phantoms. Moreover, optical properties of the top layer of a solid phantom are determined. As a numerical tool for this parameter identification, we demonstrate that the numerical scheme for the radiative transport equation (RTE) based on the method of rotated reference frames [8,9] provides an efficient numerical algorithm for SFDI partially because the method relies on the Fourier transform in the spatial-frequency domain.

Noting that the solution to the RTE is expressed as a superposition of three-dimensional singular eigenfunctions [10], we investigate the asymptotic behavior of the solution. The effect of the spatial frequency q_0 on the decay of the specific intensity is found. The longest-lived mode is controlled by q_0 and the largest Case's eigenvalue. This finding gives a theoretical reason that shallow regions can be exclusively studied by SFDI.

Then we numerically solve the RTE in the half space by the method of rotated reference frames [8]. The method of rotated reference frames in the half space was developed for a point source [9,11] and for a spatially oscillating source [12]. In [9,13], the subtraction of the ballistic term was considered. Using the method of rotated reference frames, the effect of surface scattering in SFDI was studied [14]. The inverse problem is solved by the Levenberg–Marquardt algorithm. With our approach, optical parameters of the top layer of a layered solid phantom made of epoxy resin was determined within 1 s on a laptop computer.

The remainder of this paper is organized as follows. In Section 2, we study the asymptotic behavior of the specific intensity to consider the penetration depth of near-infrared light illuminated by a spatially modulated source. In Section 3, reconstruction of optical properties of the top layer of layered media is considered for different numerical phantoms. Optical properties are estimated using a solid phantom in Section 4. Discussion and conclusions are given in Section 5. In Appendix A, two types of the diffusion approximation are introduced. Appendix B is devoted to the eigenmode expansion and numerical algorithm of the RTE.

2. DECAY BEHAVIOR

Let us consider near-infrared light propagation in the half space. Let Ω be the half space, i.e.,

$$\Omega = \left\{ \mathbf{r} \in \mathbb{R}^3; -\infty < x < \infty, -\infty < y < \infty, 0 < z < \infty \right\}.$$
(1)

Let $\partial \Omega$ be the boundary of Ω , i.e., the x - y plane. The specific intensity at position \mathbf{r} ($\mathbf{r} = (\boldsymbol{\rho}, z)^T$, $\boldsymbol{\rho} = (x, y)^T$) in direction $\hat{\mathbf{s}}$ is denoted by $I(\mathbf{r}, \hat{\mathbf{s}})$, where $\hat{\mathbf{s}}$ is a unit vector specified by the polar angle $\vartheta \in [0, \pi]$ and azimuthal angle $\varphi \in [0, 2\pi)$. Let $d\hat{\mathbf{s}}$ denote sin $\vartheta d\vartheta d\varphi$. The RTE is written as

$$\begin{cases} (\hat{\mathbf{s}} \cdot \nabla + \mu_t) I(\mathbf{r}, \hat{\mathbf{s}}) = \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \hat{\mathbf{s}}') I(\mathbf{r}, \hat{\mathbf{s}}') d\hat{\mathbf{s}}', \\ (\mathbf{r}, \hat{\mathbf{s}}) \in \Omega \times \mathbb{S}^2, \\ I(\mathbf{r}, \hat{\mathbf{s}}) = R_n(\hat{\mathbf{s}} \cdot \hat{\mathbf{z}}) I(\mathbf{r}, \hat{\mathbf{s}}_R) + I_{\text{inc}}(\mathbf{r}, \hat{\mathbf{s}}), \quad (\mathbf{r}, \hat{\mathbf{s}}) \in \Gamma_-, \end{cases}$$
(2)

where total attenuation μ_t is the sum of absorption coefficient μ_a and scattering coefficient μ_s , which are both assumed to be positive constants, and $p(\hat{\mathbf{s}}, \hat{\mathbf{s}}')$ is the scattering phase function. We assume that $p(\hat{\mathbf{s}}, \hat{\mathbf{s}}')$ is given by

$$p(\hat{\mathbf{s}}, \hat{\mathbf{s}}') = \sum_{l=0}^{l_{\max}} \sum_{m=-l}^{l} g^{l} Y_{lm}(\hat{\mathbf{s}}) Y_{lm}^{*}(\hat{\mathbf{s}}'),$$
(3)

where l_{\max} is a positive integer, $g \in (-1, 1)$ is a constant, and the superscript * denotes a complex conjugate. Spherical harmonics $Y_{lm}(\hat{\mathbf{s}})$ are defined by

$$Y_{lm}(\hat{\mathbf{s}}) = \sqrt{\frac{2l+1}{4\pi} \frac{(l-m)!}{(l+m)!}} P_l^m(\cos\vartheta) e^{im\varphi},$$
 (4)

where $P_l^m(\mu)$ are associated Legendre polynomials. Throughout the paper, we set $l_{max} = 9$. Moreover,

$$\Gamma_{-} = \left\{ (\mathbf{r}, \, \hat{\mathbf{s}}) \in \partial\Omega \times \mathbb{S}^{2}; \, \nu(\mathbf{r}) \cdot \hat{\mathbf{s}} < 0 \right\}$$
$$= \left\{ (\mathbf{r}, \, \hat{\mathbf{s}}) \in \partial\Omega \times \mathbb{S}^{2}_{+} \right\},$$
(5)

where $v(\mathbf{r})$ is the outer unit vector normal to $\mathbf{r} \in \partial \Omega$, and \mathbb{S}^2_+ denotes the set of unit vectors in inward directions. We give the incident beam $I_{\text{inc}}(\mathbf{r}, \hat{\mathbf{s}})$ as

$$I_{\rm inc}(\mathbf{r},\,\hat{\mathbf{s}}) = e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}}\delta(\hat{\mathbf{s}}-\hat{\mathbf{z}}), \quad \mathbf{q}_0 \in \mathbb{R}^2, \tag{6}$$

where $\hat{\mathbf{z}}$ is the unit vector in the positive *z* direction. The Fresnel reflection for the ratio **n** between the refractive indices inside and outside is also considered in the boundary condition. The direction $\hat{\mathbf{s}}_R$ is specified by the polar angle $\pi - \vartheta$ and azimuthal angle φ . Assuming unpolarized light, the Fresnel coefficient $R_n(\mu)$ ($0 < \mu \le 1$) is given by [15]

$$R_{\mathfrak{n}}(\mu) = \begin{cases} \frac{1}{2} \left(\left(\frac{\mu - \mathfrak{n}\mu_0}{\mu + \mathfrak{n}\mu_0} \right)^2 + \left(\frac{\mu_0 - \mathfrak{n}\mu}{\mu_0 + \mathfrak{n}\mu} \right)^2 \right) & \text{for } \mu \ge \mu_c, \\ 1 & \text{for } \mu < \mu_c, \end{cases}$$
(7)

where $\mu_0 = \sqrt{1 - n^2(1 - \mu^2)}$, and $\mu_c = \sqrt{n^2 - 1}/n$.

Suppose that \mathbf{r}_d is a point on $\partial \Omega$. We detect the hemispheric flux

$$J_{+}(\mathbf{r}_{d}) = \int_{0}^{2\pi} \int_{\pi}^{2\pi} (\cos\vartheta) I(\mathbf{r}_{d}, \hat{\mathbf{s}}) \sin\vartheta \, \mathrm{d}\vartheta \, \mathrm{d}\varphi = -A_{\mathrm{RTE}}(q_{0})e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}},$$
(8)

where $q_0 = |\mathbf{q}_0|$, and $A_{\text{RTE}}(q_0)$ is given by (B41) in Appendix B.

The fluence of the specific intensity is asymptotically governed by the diffusion equation. Let $u(\mathbf{r})$ denote the solution to the diffusion equation. To do the diffusion approximation, we split the specific intensity into two terms (see Appendix A): $I(\mathbf{r}, \hat{\mathbf{s}}) = I_0(\mathbf{r}, \hat{\mathbf{s}}) + I_1(\mathbf{r}, \hat{\mathbf{s}})$, where $I_0(\mathbf{r}, \hat{\mathbf{s}})$ satisfies

$$\begin{cases} \left(\hat{\mathbf{s}} \cdot \nabla + \bar{\mu}\right) I_0(\mathbf{r}, \, \hat{\mathbf{s}}) = 0, & (\mathbf{r}, \, \hat{\mathbf{s}}) \in \Omega \times \mathbb{S}^2, \\ I_0(\mathbf{r}, \, \hat{\mathbf{s}}) = I_{\text{inc}}(\mathbf{r}, \, \hat{\mathbf{s}}), & (\mathbf{r}, \, \hat{\mathbf{s}}) \in \Gamma_-. \end{cases}$$
(9)

Here, different choices are possible for $\bar{\mu}$ [16,17]. Then I_1 is determined depending on the choice of $\bar{\mu}$. The P_1 approximation is made for this I_1 , and eventually we arrive at the diffusion equation. We introduce

$$\mu'_{s} = (1 - g)\mu_{s}, \quad \mu_{*} = \mu_{a} + \mu'_{s}, \quad \mu_{eff} = \sqrt{3\mu_{a}\mu_{*}}.$$
 (10)

Probably the most naive choice is

$$\bar{\mu} = \mu_a + \mu_s. \tag{11}$$

We call this diffusion approximation DA1. In this case, from (A9), (A11), (A13), and (A16), we obtain

$$u(\mathbf{r}) = v_{\text{DA1}}(z)e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}},$$
(12)

where

Another choice is to set [18]

$$\bar{\mu} = \mu_*. \tag{14}$$

We refer to this diffusion approximation as DA2. In this case, from (A9), (A25), and (A26), we have

$$u(\mathbf{r}) = v_{\text{DA2}}(z)e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}},\tag{15}$$

where

$$v_{\text{DA2}}(z) = \frac{3\mu'_{s}\mu_{*}}{\mu_{*}^{2} - \mu_{\text{eff}}^{2} - q_{0}^{2}} \times \left(\frac{\mu_{*} + 3\mu_{*}/\zeta}{\sqrt{\mu_{\text{eff}}^{2} + q_{0}^{2}} + 3\mu_{*}/\zeta} e^{-\sqrt{\mu_{\text{eff}}^{2} + q_{0}^{2}} z} - e^{-\mu_{*}z}\right).$$
(16)

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Using the diffusion approximation, the detected light $J_+(\mathbf{r}_d)$ can be expressed as

$$-A_{\text{DA1}}(q_0)e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}}$$
 or $-A_{\text{DA2}}(q_0)e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}}$, (17)

depending on DA1 or DA2. They are given by (A19) or (A29) in Appendix A.

In the singular-eigenfunction approach [19], the separation constant ν is either an eigenvalue $\nu_j(M) > 1$ $(j = 1, ..., J^M)$ or in the continuous spectrum (0, 1). Although there is only one positive eigenvalue ν_0 in the case of isotropic scattering, in general there are multiple eigenvalues. We order them as $\nu_1 > \nu_2 > \cdots > \nu_{J^M} > 1$ for each M.

Case's method can be extended to three dimensions [10], and the solution to (2) can be expressed as

$$I(\mathbf{r}, \hat{\mathbf{s}}) = e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}} \sum_{M=-l_{\max}}^{l_{\max}} \\ \times \left[\sum_{j=0}^{J^{M}-1} a_{j}^{M} \Psi_{\nu_{j}(M)}^{M}(\hat{\mathbf{s}}, \mathbf{q}_{0}) e^{-\mu_{t}\hat{k}_{z}(\nu_{j}(M)q_{0})z/\nu_{j}(M)} \right. \\ \left. + \int_{0}^{1} a^{M}(\nu) \Psi_{\nu}^{M}(\hat{\mathbf{s}}, \mathbf{q}_{0}) e^{-\mu_{t}\hat{k}_{z}(\nu q_{0})z/\nu} \mathrm{d}\nu \right],$$
(18)

with coefficients a_j^M , $a^M(v)$. Here, $\Psi_{v_j(M)}^M(\hat{\mathbf{s}}, \mathbf{q}_0)$, $\Psi_v^M(\hat{\mathbf{s}}, \mathbf{q}_0)$ are three-dimensional singular eigenfunctions introduced in Appendix B.

We note that $\nu_0 = \nu_1(0) > 1$ is the largest eigenvalue. Let us define

$$I_{0}(\mathbf{r}, \hat{\mathbf{s}}) = I_{0}(\boldsymbol{\rho}, z, \hat{\mathbf{s}}) = e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}} a_{0}^{0} \Psi_{\nu_{0}}^{0}(\hat{\mathbf{s}}, \mathbf{q}_{0}) e^{-\mu_{t}\hat{k}_{z}(\nu_{0}q_{0})z/\nu_{0}}.$$
(19)

When z is large, the contribution of the mode I_0 dominates:

$$\|I(\cdot, z, \cdot) - I_0(\cdot, z, \cdot)\|_{L^{\infty}(\mathbb{R}^2; L^{\infty}(\mathbb{S}^2))}$$
$$= o\left(\exp\left(-z\sqrt{\left(\frac{\mu_t}{\nu_0}\right)^2 + q_0^2}\right)\right),$$
(20)

as $z \to \infty$.

In the case of isotropic scattering (g = 0) [19], the eigenvalue ν_0 satisfies $1 - (\mu_s/\mu_t)\nu_0 \tanh^{-1}(1/\nu_0) = 0$. When $\mu_a \ll \mu_s$ as is typical in biological tissue, we have $1/\nu_0 \approx \sqrt{3(1 - \mu_s/\mu_t)}$ and

$$\frac{\mu_t}{\nu_0} \approx \sqrt{3\mu_a \mu_t}.$$
 (21)

Therefore, we have $I \sim \exp(-z\sqrt{3\mu_a\mu_t + q_0^2})$.

In the general case of $g \neq 0$, we can estimate v_0 using the fact that Case's eigenvalues are approximately obtained as eigenvalues of a tridiagonal matrix B(M) [see (B29)]. It is found [16] that

$$\frac{\sqrt{1+\eta}}{\sqrt{3\frac{\mu_a}{\mu_t}\left(1-g\frac{\mu_s}{\mu_t}\right)}} \le \nu_0 \le \frac{1+\sqrt{\eta}}{\sqrt{3\frac{\mu_a}{\mu_t}\left(1-g\frac{\mu_s}{\mu_t}\right)}},$$
 (22)

where

$$\eta = \frac{4}{5} \frac{\mu_a}{\mu_a + \mu_s \left(1 - g^2\right)}.$$
 (23)

When μ_a is small, we have

$$\frac{\mu_t}{\nu_0} \approx \sqrt{3\mu_a(\mu_t - g\mu_s)} \approx \sqrt{3\mu_a\mu'_s},$$
 (24)

where $\mu'_s = (1 - g)\mu_s$. Thus, if μ_a is small, we have in general

$$I \sim e^{-z\sqrt{3\mu_a \mu_s' + q_0^2}}.$$
 (25)

When μ'_s is large, the asymptotic decay of the diffusion equation is given by $\exp(-z\sqrt{\mu_{\text{eff}}^2 + q_0^2})$ in both diffusion approximations given in (12) and (15). Note that $\mu_{\text{eff}} \approx \sqrt{3\mu_a \mu'_s}$ for small μ_a . Since $\mu_t < v_0 < \mu_{\text{eff}}$ if $\eta > 0$ is taken into account, the specific intensity decays slower than the prediction by the diffusion approximation.

If q_0 is large such that $\mu_t < \sqrt{\mu_{\text{eff}}^2 + q_0^2}$ or $\mu_* < \sqrt{\mu_{\text{eff}}^2 + q_0^2}$, the asymptotic decay from the diffusion approximation given in (12) or (15) becomes $\exp(-\mu_t z)$ or $\exp(-\mu_* z)$, and in either case, the asymptotic behaviors of the RTE and diffusion equation are quite different.

3. SPATIAL-FREQUENCY DOMAIN IMAGING

Although different choices are possible, we modulate the illuminating light in the x direction and give the vector \mathbf{q}_0 as

$$\mathbf{q}_0 = (2\pi f, 0)^T$$
. (26)

When the sample is illuminated by the source,

$$I_{\rm inc}(\mathbf{r},\,\hat{\mathbf{s}}) = \cos(2\pi f x)\delta(\hat{\mathbf{s}} - \hat{\mathbf{z}}).$$
(27)

The measured light is expressed as

$$J_{+}(\mathbf{r}_{d}) = -A(q_{0})\cos(2\pi f x_{d}),$$
(28)

where x_d is the first component of \mathbf{r}_d . Here, $A(q_0)$ depends on $q_0 = 2\pi f$.

We use the amplitude $A(q_0)$ to reconstruct optical properties. Let N_f be the number of spatial frequencies used for reconstruction. We have

$$f = f_1, \dots, f_{N_f}.$$
 (29)

Correspondingly, we write the forward data as $A(q_0^{(i)})$ $(i = 1, ..., N_f)$. Let $\mathbf{y} \in \mathbb{R}^{N_f}$ be a vector defined as $\mathbf{y} = (A(q_0^{(i)}))$. We can express $A_{\text{RTE}}(q_0)$ as $A_{\text{RTE}}(q_0^{(i)})$ $(i = 1, ..., N_f)$. Similarly, $A_{\text{DA1}}(q_0^{(i)})$ and $A_{\text{DA2}}(q_0^{(i)})$ are introduced. Then computed values are stored in a vector $\mathbf{F} \in \mathbb{R}^{N_f}$.

Parameters μ_a , μ'_s are determined by the Levenberg– Marquardt algorithm [20,21]. To run the inversion algorithm with scaled variables ξ_1 , ξ_2 , we express the optical properties as [22,23]

$$\mu_a = \mu_a^{(0)} e^{\xi_1}, \quad \mu_s' = \mu_s'^{(0)} e^{\xi_2},$$
 (30)

where $\mu_a^{(0)}$, $\mu_s'^{(0)}$ are initial guesses. That is, $\xi_1 = \ln(\mu_a/\mu_a^{(0)})$, $\xi_2 = \ln(\mu_s'/\mu_s'^{(0)})$.

We express $\mathbf{F} = \mathbf{F}(\boldsymbol{\xi})$, where $\boldsymbol{\xi} = (\xi_1, \xi_2)^T$. We wish to find

$$\boldsymbol{\xi}_* = \arg\min_{\boldsymbol{\xi}} \left| \mathbf{y} - \mathbf{F}(\boldsymbol{\xi}) \right|^2.$$
(31)

Let $\boldsymbol{\xi}^k$ be the estimated solution at the *k*th iteration. Starting with the initial guess $\boldsymbol{\xi}^0 = \mathbf{0}$, the solution is given by $\boldsymbol{\xi}_* = \lim_{k \to \infty} \boldsymbol{\xi}^k$. The FORTRAN library MINPACK [24] was used for the numerical calculation of the Levenberg–Marquardt method. Reconstructed values are obtained as

$$\mu_a = \mu_a^{(0)} \exp(\xi_{*,1}), \quad \mu'_s = \mu'_s(0) \exp(\xi_{*,2}).$$
 (32)

Below, we will perform several parameter identifications using numerical phantoms.

A. Thin Slabs

We perform parameter identification for numerical slab phantoms of size 90 mm × 90 mm × *L* as shown in Fig. 1. The thickness *L* changes from 1 mm to 10 mm. The optical parameters of slabs are set to $\mu_a = 0.02 \text{ mm}^{-1}$, $\mu_s = 10 \text{ mm}^{-1}$, and g = 0.9. Moreover, n = 1 (vacuum boundary condition). We suppose μ_a , μ'_s are unknown. We set $N_f = 2$, and $f_1 = 0.1$, $f_2 = 0.2 \text{ (mm}^{-1})$.

As the forward data, the hemispheric flux is computed by Monte Carlo simulations and stored in **y**. In each run, 10^8 photons are launched. The vector $\mathbf{F} = (A_{\text{RTE}}(q_0^{(i)}))$ (i = 1, 2) is computed from the RTE [see (B41)]. For the inverse problem of parameter identification, initial values are set to $(\mu_a^{(0)}, \mu_s^{(0)}) = (0.01 \text{ mm}^{-1}, 10 \text{ mm}^{-1}).$

Figure 2 shows estimated μ_a and μ'_s for the slabs. Estimated values are (μ_a mm⁻¹, μ'_s mm⁻¹) = (0.18, 1.2), (0.055, 1.1), (0.033, 1.1), (0.023, 1.0), (0.023, 1.0), (0.024, 1.1), (0.023, 1.0), (0.021, 1.0), (0.023, 1.0), and (0.024, 1.0) for L = 1 mm, 2 mm, 3 mm, 4 mm, 5 mm, 6 mm, 7 mm, 8 mm, 9 mm, and 10 mm, respectively. We see that for slabs of thickness larger than 4 mm, optical properties are correctly obtained.

The behavior in Fig. 2 is implied in the decay in (25) derived in Section 2. In the present situation, we have

$$I \sim e^{-2\pi f z}.$$
 (33)

For $f = 0.1 \text{ mm}^{-1}$, we have $e^{-2\pi f \cdot 3} = 0.15$ and $e^{-2\pi f \cdot 4} = 0.08$. Hence, the specific intensity is reduced by more than one tenth when the thickness of the slab is 4 mm or larger. Recently, an intensive study of Monte Carlo look-up tables was reported for relations between the penetration depths of photons and spatial frequencies [25]. Our conclusion in Fig. 2 is consistent with their results.



Fig. 1. Thin slab.



Fig. 2. Upper panel shows estimated $\mu_a \text{ mm}^{-1}$ for slabs of thicknesses 1 mm through 10 mm. The lower panel shows estimated $\mu'_{c} \text{ mm}^{-1}$ for the same slabs. Dotted lines show true values.



Table 1. Reconstructed μ_a and μ'_s (mm⁻¹) of the Top Layer of the Two-Layer Medium by RTE for $\mu_a = 0.01 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathrm{mm}^{-1})$	Error
True	0.02	_	1.0	
RTE	0.0079	0.60	1.1	0.14

B. Two-Layer Media

Let us consider two-layer media, which have top and bottom layers. As shown in Fig. 3, the thickness of the top layer is 6 mm, and the bottom layer has a thickness of 30 mm, which can be regarded as a semi-infinite medium. We set $N_f = 2$, and $f_1 = 0.1$, $f_2 = 0.2$ (mm⁻¹). The scattering coefficient and anisotropic factor are fixed to $\mu_s = 10 \text{ mm}^{-1}$ and g = 0.9, respectively, in the entire medium. The refractive index is set to $\mathbf{n} = 1.4$. The absorption coefficient in the top layer is $\mu_a = 0.02 \text{ mm}^{-1}$. The absorption coefficient μ_a in the bottom layer takes values 0.01 mm^{-1} and 0.03 mm^{-1} . We use $A_{\text{RTE}}(q_0^{(i)})$ (i = 1, 2) for reconstruction. The forward data were computed by Monte Carlo simulations. Initial values were set to $(\mu_a^{(0)}, \mu_s^{(0)}) = (0.01 \text{ mm}^{-1}, 10 \text{ mm}^{-1})$.

Below, we present obtained μ_a , μ'_s in the top layer when the absorption coefficient of the bottom layer is 0.01 mm⁻¹ (Table 1) and 0.03 mm⁻¹ (Table 2). To see how the estimated value is close to the true value, in Tables 1 and 2, we also give the relative error, which is defined as

Table 2. Reconstructed μ_a and μ'_s (mm⁻¹) of the Top Layer of the Two-Layer Medium by RTE for $\mu_a = 0.03 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathbf{mm}^{-1})$	Error	$\mu'_s(\mathbf{mm}^{-1})$	Error
True	0.02	_	1.0	_
RTE	0.0087	0.57	1.1	0.13
Fig. 4. Thr	ree-layer medium			0mm 2mm 9mm

|(estimated value – true value)/true value|. Numerical results show the reconstructed values are not affected by optical properties of the bottom layer.

C. Three-Layer Media

Here, we consider more complex media that have more than two layers [26,27]. We reconstruct the optical properties of the top layer of a three-layer medium, shown in Fig. 4. Their optical properties are summarized in Table 3. The depths of layers are, from the top, 10 mm, 2 mm, and 28 mm. From the top, $\mu_s = 18$, 3, and 21 (mm⁻¹). In the second layer, $\mu_a = 0.004 \text{ mm}^{-1}$. The absorption coefficients of the first and third layers vary from 0.01 mm⁻¹ to 0.03 mm⁻¹. In addition, g = 0.9 and n = 1.4. We set $N_f = 2$, and $f_1 = 1/15$, $f_2 = 1/10$ (mm⁻¹). Starting the Levenberg–Marquardt algorithm with initial values ($\mu_a^{(0)}, \mu_s^{(0)}$) = (0.01 mm⁻¹, 10 mm⁻¹), we obtain μ_a, μ'_s after about 10 iterations.

For comparison, reconstructions by the diffusion approximation are also obtained. That is, A_{DA1} in (A19) and A_{DA2} in (A29) are used in addition to A_{RTE} in (8). The forward data were prepared by Monte Carlo simulations.

Tables 4–12 show reconstructed μ_a , μ'_s when the true values in the top layer are $\mu_a = 0.01 \text{ mm}^{-1}$, 0.02 mm^{-1} , and 0.03 mm^{-1} . In each table, relative errors are also shown. The true value of μ'_s in the top layer is fixed to 1.8 mm⁻¹, as shown in Table 3. We see that reconstructed values by DA1 are closer to those by RTE compared with reconstructed values by DA2. The absorption coefficient of the bottom layer is $\mu_a = 0.01 \text{ mm}^{-1}$ in Tables 4, 7, 10, $\mu_a = 0.02 \text{ mm}^{-1}$ in Tables 5, 8, 11, and $\mu_a = 0.03 \text{ mm}^{-1}$ in Tables 6, 9, 12. In all cases, reconstructed

Table 3. Three-Layer Model^a

	$\mu_a(\mathrm{mm}^{-1})$	$\mu_s(\mathrm{mm}^{-1})$	<i>d</i> (mm)
1st layer	$\mu_{a,1}$	18	10
2nd layer	0.004	3	2
3rd layer	$\mu_{a,3}$	21	28

'In the top layer, $\mu_{a,1} = 0.01$, 0.02, or 0.03 (mm⁻¹). In the bottom layer, $\mu_{a,3} = 0.01$, 0.02, or 0.03 (mm⁻¹).

Table 4. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.01 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.01 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathrm{mm}^{-1})$	Error
True	0.01		1.8	
RTE	0.0084	0.16	1.8	0.021
DA1	0.0060	0.40	1.8	0.018
DA2	0.015	0.53	1.5	0.18

Table 5. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.01 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.02 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu'_s(\mathbf{mm}^{-1})$	Error
True	0.01	_	1.8	
RTE	0.0085	0.15	1.8	0.021
DA1	0.0060	0.40	1.8	0.018
DA2	0.015	0.53	1.5	0.18

Table 6. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.01 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.03 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathbf{m}\mathbf{m}^{-1})$	Error	$\mu_s'(\mathrm{mm}^{-1})$	Error
True	0.01		1.8	
RTE	0.0085	0.15	1.8	0.022
DA1	0.0060	0.40	1.8	0.018
DA2	0.015	0.54	1.5	0.18

Table 7. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.02 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.01 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathbf{mm}^{-1})$	Error
True	0.02		1.8	_
RTE	0.016	0.20	1.8	0.0015
DA1	0.012	0.39	1.9	0.053
DA2	0.026	0.31	1.5	0.19

Table 8. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.02 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.02 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathbf{mm}^{-1})$	Error	$\mu_s'(\mathbf{mm^{-1}})$	Error
True	0.02		1.8	
RTE	0.016	0.20	1.8	0.0013
DA1	0.012	0.39	1.9	0.053
DA2	0.026	0.31	1.5	0.19

values are not affected by the third layer.

4. SOLID PHANTOM

Figure 5(a) shows a solid phantom made of epoxy resin. The refractive index of the phantom is 1.58. The phantom has a fourlayer structure, and the width of the top layer is about 4 mm. When the phantom was made, optical properties of the top



Fig. 5. (a) Four-layer solid phantom. The width of the top layer is about 4 mm, and the width of the bottom layer is about 47 mm. Measurement setup: (b) top view and (c) side view. DMD, digital micro-mirror device. BPF, band-pass filter.

Table 9.	Reconstructed μ_a and μ'_s (mm ⁻¹) by RTE,
DA1, and	DA2 When $\mu_a = 0.02 \mathrm{mm^{-1}}$ in the Top Layer
and $\mu_a = 0$).03 mm ⁻¹ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathrm{mm}^{-1})$	Error
True	0.02		1.8	_
RTE	0.016	0.20	1.8	0.0012
DA1	0.012	0.39	1.9	0.053
DA2	0.026	0.32	1.5	0.19

Table 10. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.03 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.01 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathrm{mm}^{-1})$	Error
True	0.03	_	1.8	_
RTE	0.023	0.25	1.8	0.026
DA1	0.017	0.42	2.0	0.088
DA2	0.036	0.21	1.5	0.18

Table 11. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.03 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.02 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathbf{mm}^{-1})$	Error
True	0.03	_	1.8	_
RTE	0.023	0.25	1.8	0.026
DA1	0.017	0.42	2.0	0.088
DA2	0.036	0.21	1.5	0.18

Table 12. Reconstructed μ_a and μ'_s (mm⁻¹) by RTE, DA1, and DA2 When $\mu_a = 0.03 \text{ mm}^{-1}$ in the Top Layer and $\mu_a = 0.03 \text{ mm}^{-1}$ in the Bottom Layer

	$\mu_a(\mathrm{mm}^{-1})$	Error	$\mu_s'(\mathbf{mm}^{-1})$	Error
True	0.03	_	1.8	
RTE	0.023	0.25	1.8	0.026
DA1	0.017	0.42	2.0	0.088
DA2	0.036	0.21	1.5	0.18

layer were aimed at $\mu_a = 0.0165 \text{ mm}^{-1}$ and $\mu'_s = 1.3 \text{ mm}^{-1}$. The setup of the measurement system is shown in Figs. 5(b) and 5(c). As a near-infrared light source, a broadband halogen fiber optic illuminator (Thorlabs, OSL2) with an enhanced infrared replacement bulb (Thorlabs, OSL2BIR) was used. The projected patterns were displayed on a digital micromirror device (DMD) module (Keynote Photonics, LC4500 NIR controller). The image was formed on the sample plane through a visible-near-infrared lens (Schneider, large format F-mount lens, focal length of 28 mm, F/2.8). To capture reflection images, a camera with enhanced near-infrared sensitivity (Ximea, MQ-13RG-E2, 1280x1024 pixels, monochrome) with a visible-near-infrared lens (Edmund Optics, C series VIS-NIR fixed focal length lens, focal length of 16 mm, F/1.6) was used. The polarizer and analyzer placed in the crossed nicols configuration were inserted into the measurement system to remove the specular reflection from the sample. A band-pass filter (Edmund Optics, hard coated OD 4.0 25 nm band-pass filter, center wavelength of 800 nm, FWHM of 25 nm) was placed in front of the camera to extract the 800 nm wavelength. The top layer of the phantom was illuminated by the spatially modulated light, and the reflected light was detected by the camera.

The source term in experiments is given by

$$I_{\rm inc}(\mathbf{r},\,\hat{\mathbf{s}}) = \frac{S_0}{2} \left[1 + \cos(2\pi f x + \alpha)\right] \delta(\hat{\mathbf{s}} - \hat{\mathbf{z}}), \qquad (34)$$

where $\alpha = 2\pi p/3$ (p = 0, 1, 2). We set $N_f = 2$, and $f_1 = 0.1$, $f_2 = 0.2 \,(\text{mm}^{-1})$. We have

$$J_{+}(\mathbf{r}_{d}; \alpha) = \frac{S_{0}}{2} \left(M_{\rm DC} + M_{\rm AC} \cos(2\pi f x_{d} + \alpha) \right), \quad (35)$$

where $M_{\rm DC}$, $M_{\rm AC}$ depend on \mathbf{r}_d and f in general. Let us write $J^{(p)}_+ = J_+(\mathbf{r}_d; 2\pi p/3)$. By a straightforward calculation, we have

$$(M_{\rm AC})^2 = \frac{2}{9} \times \left[\left(J_+^{(0)} - J_+^{(1)} \right)^2 + \left(J_+^{(1)} - J_+^{(2)} \right)^2 + \left(J_+^{(2)} - J_+^{(0)} \right)^2 \right].$$
(36)

Moreover, we can write [7]

$$M_{\rm AC} = M^{\rm exp}(\mathbf{r}_d, f) A^{\rm exp}(q_0^{(i)}),$$
 (37)

where $M^{\text{exp}}(>0)$ is a constant determined by the optical system, and $q_0^{(i)} = f_i$ (i = 1, 2).

Let us consider $M_{AC,ref}$ of a reference medium, for which $A_{ref}^{exp}(q_0)$ can be numerically computed. Then we have [7]

$$A^{\exp}(q_0) = \left\langle \frac{M_{\rm AC}}{M_{\rm AC,ref}} \right\rangle A^{\exp}_{\rm ref}(q_0), \qquad (38)$$

where $\langle \cdot \rangle$ means the average in space.

We use the bottom layer of the solid phantom as the reference medium. Since the fourth layer of the phantom has the thickness 47 mm, it can be regarded as the half space even for time-resolved measurements in which two optical fibers are vertically attached on the bottom side of the phantom. From time-resolved measurements by TRS-80 (Hamamatsu Photonics), we found $\mu_a = 0.010 \text{ mm}^{-1}$ and $\mu'_s = 1.4 \text{ mm}^{-1}$ for the bottom layer. With these optical parameters, $A_{\text{ref}}^{\exp}(q_0^{(i)})$ can be computed from the RTE as $A_{\text{RTE}}(q_0^{(i)})$. Thus, $A^{\exp}(q_0^{(i)})$ in (38) are prepared and stored in a vector $\mathbf{y} \in \mathbb{R}^{N_f}$.

The Levenberg–Marquardt algorithm was run with initial values $\mu_a^{(0)} = 0.01 \text{ mm}^{-1}$ and $\mu_s^{(0)} = 10 \text{ mm}^{-1}$. We obtain

$$\mu_a = 0.016 \text{ mm}^{-1}, \quad \mu'_s = 1.0 \text{ mm}^{-1}.$$
 (39)

Other choices of initial guess, for example, $\mu_a^{(0)} = 0.02 \text{ mm}^{-1}$ and $\mu_s^{\prime (0)} = 2 \text{ mm}^{-1}$, give the same μ_a , μ_s^{\prime} given in (39). The computation time was less than 1 s with a laptop computer (MacBook Pro with 2.3 GHz Intel Core i5 and 8 GB memory).

5. DISCUSSION AND CONCLUSION

Taking advantage of the fact that near-infrared light decays rapidly for nonzero spatial frequencies, in this paper, we estimated optical properties of the top layer of the layered phantom. Indeed, SFDI has been used for the parameter identification of the top layer. Various numerical experiments developed in the present paper confirm that results by SFDI are not affected by deeper layers. Figure 2 indicates that the necessary width of the top layer is 4 mm (when $1/\mu'_s = 1$ mm). Our investigation in the asymptotic limit provides a theoretical reason for the results of numerical experiments. Using the numerical algorithm based on the method of rotated reference frames, optical properties of the top layer of a layered solid phantom were also determined.

Our numerical results show that μ'_s is more accurately reconstructed than μ_a . One of the reasons for this inaccuracy is the ill-posedness of the parameter identification. Future study on regularization might improve the results.

When two optical fibers are attached to biological tissue in the direction perpendicular to its surface with a separation of a few centimeters, the detected reflected near-infrared light contains photons from depths more than a centimeter. Although this very feature makes it possible to study the function of the human brain through the neurovascular coupling [28], the detected light is affected by optical properties of different layers. In particular, the signal from the brain is affected by skin blood flow in the scalp [29]. In this conventional way, it is not possible to extract only information on shallow regions. Photons that travel deep inside biological tissue can be excluded if the separation of two optical fibers is reduced. However, then measurements have to be conducted in a tiny space, and other difficulties related to measurements arise (see [30] and references therein). The SFDI measurement setup described in Fig. 5(a) is free from such difficulties.

For our approach to work, the top layer of a layered random medium has to be regarded as a semi-infinite medium. The necessary width of the top layer depends on spatial frequency. This can be checked by a test parameter identification for numerical phantoms.

When μ_a is not small, the decay of the specific intensity deviates from the diffusive decay and is given by (20) because then μ_t/ν_0 cannot be approximated by $\sqrt{3\mu_a\mu_s}$. Moreover if the depth is not large compared to ν_0/μ_t or $1/q_0$, not only I_0 but other modes contribute, and the decay is given by the superposition of different decays shown in (18).

APPENDIX A: DIFFUSION APPROXIMATION

We begin by decomposing *I* into the following two terms:

$$I(\mathbf{r}, \hat{\mathbf{s}}) = I_b(\mathbf{r}, \hat{\mathbf{s}}) + I_s(\mathbf{r}, \hat{\mathbf{s}}).$$
(A1)

The ballistic term I_b and scattering term I_s satisfy

$$\begin{cases} \left(\hat{\mathbf{s}} \cdot \nabla + \mu_a + \mu_s\right) I_b(\mathbf{r}, \, \hat{\mathbf{s}}) = 0, & (\mathbf{r}, \, \hat{\mathbf{s}}) \in \Omega \times \mathbb{S}^2, \\ I_b(\mathbf{r}, \, \hat{\mathbf{s}}) = I_{\text{inc}}(\mathbf{r}, \, \hat{\mathbf{s}}), & (\mathbf{r}, \, \hat{\mathbf{s}}) \in \Gamma_- \end{cases}$$
(A2)

and

$$\begin{cases} \left(\hat{\mathbf{s}} \cdot \nabla + \mu_a + \mu_s\right) I_s(\mathbf{r}, \,\hat{\mathbf{s}}) = \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \,\hat{\mathbf{s}}') I_s(\mathbf{r}, \,\hat{\mathbf{s}}') \mathrm{d}\hat{\mathbf{s}}' \\ + S(\mathbf{r}, \,\hat{\mathbf{s}}), \quad (\mathbf{r}, \,\hat{\mathbf{s}}) \in \Omega \times \mathbb{S}^2, \\ I_s(\mathbf{r}, \,\hat{\mathbf{s}}) = R_n(\hat{\mathbf{s}} \cdot \hat{\mathbf{z}}) I_s(\mathbf{r}, \,\hat{\mathbf{s}}_R), \quad (\mathbf{r}, \,\hat{\mathbf{s}}) \in \Gamma_-. \end{cases}$$
(A3)

Here, the source term for I_s is given by

$$S(\mathbf{r}, \hat{\mathbf{s}}) = \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \hat{\mathbf{s}}') I_b(\mathbf{r}, \hat{\mathbf{s}}') \mathrm{d}\hat{\mathbf{s}}'.$$
(A4)

Since

$$I_b(\mathbf{r}, \,\hat{\mathbf{s}}) = e^{-\mu_t z} e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}} \delta(\hat{\mathbf{s}} - \hat{\mathbf{z}}), \tag{A5}$$

we have

$$S(\mathbf{r}, \hat{\mathbf{s}}) = \mu_s e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}} e^{-\mu_t z} p(\hat{\mathbf{s}}, \hat{\mathbf{z}}).$$
 (A6)

Suppose that I_s weakly depends on \hat{s} and can be written as

$$I_{s}(\mathbf{r},\,\hat{\mathbf{s}}) = \frac{1}{4\pi}u(\mathbf{r}) + \frac{3}{4\pi}\mathbf{J}(\mathbf{r})\cdot\hat{\mathbf{s}},\qquad (A7)$$

where

$$u(\mathbf{r}) = \int_{\mathbb{S}^2} I_s(\mathbf{r}, \hat{\mathbf{s}}) \mathrm{d}\hat{\mathbf{s}}, \quad \mathbf{J}(\mathbf{r}) = \int_{\mathbb{S}^2} \hat{\mathbf{s}} I_s(\mathbf{r}, \hat{\mathbf{s}}) \mathrm{d}\hat{\mathbf{s}}.$$
(A8)

Let us write

$$u(\mathbf{r}) = v_{\text{DA1}}(z)e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}}.$$
 (A9)

Then from (A3), we obtain

$$\frac{\partial^2}{\partial z^2} v_{\text{DA1}} - \left(\mu_{\text{eff}}^2 + q_0^2\right) v_{\text{DA1}} = -Be^{-\mu_t z},$$
 (A10)

where

$$B = 3\mu_*\mu_s \left(1 + g\frac{\mu_t}{\mu_*}\right).$$
 (A11)

We note that

$$\mathbf{J}(\mathbf{r}) = -\frac{1}{3\mu_*} \nabla u(\mathbf{r}) + \frac{g\mu_s}{\mu_*} e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}} e^{-\mu_t z} \hat{\mathbf{z}}.$$
 (A12)

Weobtain

$$v_{\rm DA1}(z) = \frac{-B}{\mu_t^2 - \mu_{\rm eff}^2 - q_0^2} e^{-\mu_t z} + C_1 e^{-\sqrt{\mu_{\rm eff}^2 + q_0^2} z},$$
 (A13)

where C_1 is determined from the boundary condition. If we assume the diffuse boundary condition such that

$$-D_0\frac{\partial u}{\partial z} + \frac{1}{\zeta}u = 0,$$
 (A14)

where $D_0 = 1/(3\mu_*)$, and $\zeta = 2(1 + r_d)/(1 - r_d)$, with

$$r_d = -1.4399 \mathfrak{n}^{-2} + 0.7099 \mathfrak{n}^{-1} + 0.6681 + 0.0636 \mathfrak{n},$$
(A15)

we obtain

$$C_1 = \frac{B}{\mu_t^2 - \mu_{\text{eff}}^2 - q_0^2} \frac{\zeta D_0 \mu_t + 1}{\zeta D_0 \sqrt{\mu_{\text{eff}}^2 + q_0^2} + 1}.$$
 (A16)

On the boundary at \mathbf{r}_d , we have

$$u(\mathbf{r}_{d}) = \frac{3\mu_{*}\mu_{s}(1 + g\mu_{t}/\mu_{*})e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}}}{\left(\sqrt{\mu_{\text{eff}}^{2} + q_{0}^{2}} + \mu_{t}\right)\left(\sqrt{\mu_{\text{eff}}^{2} + q_{0}^{2}} + 3\mu_{*}/\zeta\right)}.$$
 (A17)

We obtain

$$J_{+}(\mathbf{r}_{d}) = \int_{\mathbb{S}^{2}_{-}} (\cos \vartheta) \left(\frac{1}{4\pi} u(\mathbf{r}_{d}) + \frac{3}{4\pi} \mathbf{J}(\mathbf{r}_{d}) \cdot \hat{\mathbf{s}} \right) d\hat{\mathbf{s}}$$
$$= -\frac{1}{4} u(\mathbf{r}_{d}) + \frac{1}{2} J_{z}(\mathbf{r}_{d})$$
$$= -A_{\mathrm{DA1}}(q_{0}) e^{i\mathbf{q}_{0} \cdot \boldsymbol{\rho}}, \qquad (A18)$$

where $u(\mathbf{r}_d)$ is given in (A17). Here, we introduced

$$A_{\rm DA1}(q_0) = \left(\frac{1}{4} + \frac{1}{2\zeta}\right) v_{\rm DA1}(0) - \frac{g\mu_s}{2\mu_*}.$$
 (A19)

Next, instead of (A2), let us introduce the ballistic term as $(\hat{\mathbf{s}} \cdot \nabla + \mu_*)I_0 = 0$; we obtain

$$I_0(\mathbf{r}, \,\hat{\mathbf{s}}) = e^{-\mu_* z} e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}} \delta(\hat{\mathbf{s}} - \hat{\mathbf{z}}), \qquad (A20)$$

and

$$S(\mathbf{r}, \hat{\mathbf{s}}) = (\mu_* - \mu_t) I_0(\mathbf{r}, \hat{\mathbf{s}}) + \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \hat{\mathbf{s}}') I_0(\mathbf{r}, \hat{\mathbf{s}}') d\hat{\mathbf{s}}'$$
$$= \left(-\mu_s g\delta(\hat{\mathbf{s}} - \hat{\mathbf{z}}) + \mu_s p(\hat{\mathbf{s}}, \hat{\mathbf{z}})\right) e^{-\mu_* z} e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}}.$$
(A21)

We can write

$$u(\mathbf{r}) = v_{\text{DA2}}(z)e^{i\mathbf{q}_0\cdot\boldsymbol{\rho}}.$$
 (A22)

After similar calculations, we obtain

$$\mathbf{J}(\mathbf{r}) = -D_0 \nabla u(\mathbf{r}), \qquad (A23)$$

and

$$\frac{\partial^2}{\partial z^2} v_{\text{DA2}} - \left(\mu_{\text{eff}}^2 + q_0^2\right) v_{\text{DA2}} = -3\mu_* \mu_s' e^{-\mu_* z}.$$
 (A24)

Hence,

$$v_{\text{DA2}}(z) = \frac{-3\mu_*\mu'_s}{\mu_*^2 - \mu_{\text{eff}}^2 - q_0^2} e^{-\mu_* z} + C_2 e^{-\sqrt{\mu_{\text{eff}}^2 + q_0^2} z}, \quad \text{(A25)}$$

where

$$C_2 = \frac{3\mu_*\mu'_s}{\mu^2_* - \mu^2_{\text{eff}} - q_0^2} \frac{\zeta D_0 \mu_* + 1}{\zeta D_0 \sqrt{\mu^2_{\text{eff}} + q_0^2} + 1}.$$
 (A26)

Weobtain

$$u(\mathbf{r}_{d}) = \frac{3\mu_{*}\mu_{s}'e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}}}{\left(\sqrt{\mu_{\text{eff}}^{2} + q_{0}^{2}} + \mu_{*}\right)\left(\sqrt{\mu_{\text{eff}}^{2} + q_{0}^{2}} + 3\mu_{*}/\zeta\right)}.$$
(A27)

In this case, we obtain

$$J_{+}(\mathbf{r}_{d}) = -A_{\text{DA2}}(q_{0})e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}},$$
 (A28)

where

$$A_{\text{DA2}}(q_0) = \left(\frac{1}{4} + \frac{1}{2\zeta}\right) v_{\text{DA2}}(0).$$
 (A29)

The difference between (A17) and (A27) is small when g is small. In this paper, we use (A17), for which the decomposition is compatible with the ballistic subtraction developed in Appendix B.

APPENDIX B: THE METHOD OF ROTATED REFERENCE FRAMES

For the method of rotated reference frames in the half space, the subtraction of the ballistic term was considered [13]. Here, we will compute the hemispheric flux J_+ following [13] with a spatially oscillating source term.

1. PRELIMINARY

We begin with the one-dimensional RTE:

$$\begin{cases} \left(\cos\vartheta \frac{\partial}{\partial z} + \mu_t\right) I_1(z, \hat{\mathbf{s}}) = \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \hat{\mathbf{s}}') I_1(z, \hat{\mathbf{s}}') \mathrm{d}\hat{\mathbf{s}}', \\ (z, \hat{\mathbf{s}}) \in (0, \infty) \times \mathbb{S}^2, \\ I_1(z, \hat{\mathbf{s}}) = R_n(\mu) I_1(z, \hat{\mathbf{s}}_R) + g_1(z, \hat{\mathbf{s}}), \\ \hat{\mathbf{s}} \in \mathbb{S}^2_+ \quad \text{at } z = 0 \end{cases}$$
(B1)

with a source term $g_1(z, \hat{s})$. In one-dimensional transport theory, it is known that the solution I_1 is expressed as [19]

$$I_{1}(z, \hat{\mathbf{s}}) = \sum_{M=-l_{\max}}^{l_{\max}} \left[\sum_{j=1}^{J^{M}} \tilde{a}_{j}^{M} \Phi_{\nu_{j}(M)}^{M}(\hat{\mathbf{s}}) e^{-\mu_{t} z/\nu_{j}(M)} + \int_{0}^{1} \tilde{a}^{M}(\nu) \Phi_{\nu}^{M}(\hat{\mathbf{s}}) e^{-\mu_{t} z/\nu} d\nu \right],$$
(B2)

where coefficients \tilde{a}_j^M , $\tilde{a}^M(v)$ are determined from the boundary condition. Here, $\Phi_v^M(\hat{s})$ ($v = v_j(M)$ or $v \in (0, 1)$) is given by [19,31–33]

$$\Phi_{\nu}^{M}(\hat{\mathbf{s}}) = \phi^{M}(\nu, \cos\vartheta) \left(1 - \cos^{2}\vartheta\right)^{|M|/2} e^{iM\varphi}, \qquad (B3)$$

where $\phi^M(v, \cos \vartheta)$ is called Case's singular eigenfunction. Let us introduce $\sigma_l > 0$ as

$$\sigma_l = \mu_t - \mu_s g^l = \mu_a + (1 - g^l) \mu_s$$
 (B4)

and introduce $b_l(m)$ as

$$b_l(m) = \sqrt{(l^2 - m^2)/((4l^2 - 1)\sigma_l\sigma_{l-1})}.$$
 (B5)

We consider the normalized Chandrasekhar polynomial $g_l^m(x)$, which satisfies the following three-term recurrence relation [34,35]:

$$\frac{x}{\mu_{l}}\sqrt{(2l+1)\sigma_{l}}g_{l}^{m}(x) = b_{l+1}(m)\sqrt{(2l+3)\sigma_{l+1}}g_{l+1}^{m}(x) + b_{l}(m)\sqrt{(2l-1)\sigma_{l-1}}g_{l-1}^{m}(x)$$
(B6)

for l > m and $m \ge 0$. We have

$$g_m^m(x) = \frac{(2m-1)!!}{\sqrt{(2m)!}} = \frac{\sqrt{(2m)!}}{2^m m!},$$
$$g_l^{-m}(x) = (-1)^m g_l^m(x),$$
$$g_l^m(-x) = (-1)^{l+m} g_l^m(x).$$
(B7)

Now, eigenvalues $v_j(M)$ are zeros of g_l^M as $l \to \infty$ [36]. In (B28), we numerically obtain $v_j(M)$ and $v \in (0, 1)$ as eigenvalues of a tridiagonal matrix.

Let $f(\hat{\mathbf{s}})$ be a function of $\hat{\mathbf{s}} \in \mathbb{S}^2$, which can be expressed as $f(\hat{\mathbf{s}}) = \sum_{l=0}^{\infty} \sum_{m=-l}^{l} f_{lm} Y_{lm}(\hat{\mathbf{s}})$, with coefficients f_{lm} . We introduce $\mathcal{R}_{\hat{\mathbf{k}}}$ for a unit vector $\hat{\mathbf{k}} \in \mathbb{C}$ ($\hat{\mathbf{k}} \cdot \hat{\mathbf{k}} = 1$) as

$$\mathcal{R}_{\hat{\mathbf{k}}} f(\hat{\mathbf{s}}) = \sum_{l=0}^{\infty} \sum_{m=-l}^{l} f_{lm} Y_{lm}(\hat{\mathbf{s}}; \hat{\mathbf{k}}),$$
(B8)

where [8]

$$Y_{lm}(\hat{\mathbf{s}}; \,\hat{\mathbf{k}}) = \sum_{m'=-l}^{l} e^{-im'\varphi_{\hat{\mathbf{k}}}} d_{m'm}^{l}(\vartheta_{\hat{\mathbf{k}}}) Y_{lm'}(\hat{\mathbf{s}}).$$
(B9)

Here, $\varphi_{\hat{\mathbf{k}}}$, $\vartheta_{\hat{\mathbf{k}}}$ are the azimuthal and polar angles of $\hat{\mathbf{k}}$, and $d_{m'm}^l$ are the Wigner d matrices. We choose the branch cut of the square root function from zero to ∞ , so $0 \leq \arg(\sqrt{z}) < \pi$ for arbitrary $z \in \mathbb{C}$. That is, by $\mathcal{R}_{\hat{\mathbf{k}}}$, we measure angles in $f(\hat{\mathbf{s}})$ in the reference frame rotated so that the z axis lies in the direction of $\hat{\mathbf{k}}$.

Indeed, ν is either an eigenvalue or in the continuous spectrum. We take the specific form of $\hat{\mathbf{k}} = \hat{\mathbf{k}}(\nu, \mathbf{q})$ given below, which depends on ν and $\mathbf{q} \in \mathbb{R}^2$:

$$\hat{\mathbf{k}}(v, \mathbf{q}) = \left(-iv\frac{\mathbf{q}}{\mu_t}, \ \hat{\mathbf{k}}_z(vq)\right), \quad \hat{k}_z(vq) = \sqrt{1 + \left(vq/\mu_t\right)^2},$$
(B10)

where $q = |\mathbf{q}|$. We note that $\hat{\mathbf{k}}(\nu, \mathbf{q}) \cdot \hat{\mathbf{k}}(\nu, \mathbf{q}) = 1$. We obtain

$$\varphi_{\hat{\mathbf{k}}} = \begin{cases} \varphi_{\mathbf{q}} + \pi & \text{for } \nu > 0, \\ \varphi_{\mathbf{q}} & \text{for } \nu < 0, \end{cases}$$
(B11)

where $\varphi_{\mathbf{q}}$ is the polar angle of \mathbf{q} , and

$$\cos\vartheta_{\hat{\mathbf{k}}} = \hat{\mathbf{k}} \cdot \hat{\mathbf{z}} = \hat{k}_z, \quad \sin\vartheta_{\hat{\mathbf{k}}} = \sqrt{1 - \cos^2\vartheta_{\hat{\mathbf{k}}}} = i|\nu q|. \quad \textbf{(B12)}$$

We note that $d_{m'm}^l(\vartheta_{\hat{\mathbf{k}}})$ depends on q but is independent of $\varphi_{\mathbf{q}}$. Hence, we write

$$d_{m'm}^{l}(\vartheta_{\hat{\mathbf{k}}}) = d_{m'm}^{l}[i\tau(\nu q)].$$
 (B13)

We define

$$\Psi^{M}_{\nu_{j}(M)}(\hat{\mathbf{s}}, \mathbf{q}_{0}) = \mathcal{R}_{\hat{\mathbf{k}}(\nu_{j}(M), \mathbf{q}_{0})} \Phi^{M}_{\nu_{j}(M)}(\hat{\mathbf{s}}),$$

$$\Psi^{M}_{\nu}(\hat{\mathbf{s}}, \mathbf{q}_{0}) = \mathcal{R}_{\hat{\mathbf{k}}(\nu, \mathbf{q}_{0})} \Phi^{M}_{\nu}(\hat{\mathbf{s}}).$$
 (B14)

2. BALLISTIC SUBTRACTION

As was done in Appendix A, we consider I_s by subtracting I_b from *I*. Let us introduce the particular solution I_p as

$$(\hat{\mathbf{s}} \cdot \nabla + \mu_a + \mu_s) I_p(\mathbf{r}, \, \hat{\mathbf{s}}) = \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \, \hat{\mathbf{s}}') I_p(\mathbf{r}, \, \hat{\mathbf{s}}') \mathrm{d}\hat{\mathbf{s}}'$$

+ $\Theta(z) S(\mathbf{r}, \, \hat{\mathbf{s}}), \, (\mathbf{r}, \, \hat{\mathbf{s}}) \in \mathbb{R}^3 \times \mathbb{S}^2,$
(B15)

where $\Theta(\cdot)$ is the step function. Then we can calculate I_s as

$$I_s = I_p + \psi, \tag{B16}$$

where ψ satisfies

$$\begin{aligned} \left(\hat{\mathbf{s}} \cdot \nabla + \mu_a + \mu_s \right) \psi(\mathbf{r}, \, \hat{\mathbf{s}}) &= \mu_s \int_{\mathbb{S}^2} p(\hat{\mathbf{s}}, \, \hat{\mathbf{s}}') \psi(\mathbf{r}, \, \hat{\mathbf{s}}') d\hat{\mathbf{s}}', \\ (\mathbf{r}, \, \hat{\mathbf{s}}) &\in \Omega \times \mathbb{S}^2, \\ \psi(\mathbf{r}, \, \hat{\mathbf{s}}) &= R_n(\hat{\mathbf{s}} \cdot \hat{\mathbf{z}}) \psi(\mathbf{r}, \, \hat{\mathbf{s}}_R) + I_{\text{inc}}^{(p)}(\mathbf{r}, \, \hat{\mathbf{s}}), \\ (\mathbf{r}, \, \hat{\mathbf{s}}) &\in \Gamma_- \\ (\mathbf{B17}) \end{aligned}$$

with

$$I_{\rm inc}^{(p)}(\mathbf{r}, \hat{\mathbf{s}}) = R_{\mathfrak{n}}(\hat{\mathbf{s}} \cdot \hat{\mathbf{z}}) I_{p}(\mathbf{r}, \hat{\mathbf{s}}_{R}) - I_{p}(\mathbf{r}, \hat{\mathbf{s}}), \quad r \in \partial \Omega.$$
(B18)

3. PARTICULAR SOLUTION

To find $I_p(\mathbf{r}, \hat{\mathbf{s}})$, we write

$$I_{p}(\mathbf{r}, \hat{\mathbf{s}}) = \mu_{s} e^{i\mathbf{q}_{0}\cdot\boldsymbol{\rho}} e^{-\mu_{t}z} \Theta(z) \sum_{l=0}^{l_{\max}} \sum_{m=-l}^{l} \eta_{lm} e^{-im\varphi_{\mathbf{q}_{0}}} Y_{lm}(\hat{\mathbf{s}}).$$
(B19)

By multiplying $Y_{lm}^*(\hat{\mathbf{s}})$ on both sides of the RTE for I_p and integrating over $\hat{\mathbf{s}} \in \mathbb{S}^2$, we arrive at the following linear system that determines η_{lm} :

$$\sum_{l'm'} \left[\frac{iq_0}{2} \left(-\delta_{m',m-1} \delta_{l',l-1} \sqrt{(l+m)(l+m-1)} + \delta_{m',m-1} \delta_{l',l+1} \sqrt{(l'-m')(l'-m'-1)} + \delta_{m',m+1} \delta_{l',l-1} \sqrt{(l-m)(l-m-1)} - \delta_{m',m+1} \delta_{l',l-1} \sqrt{(l'+m')(l'+m'-1)} \right) - \mu_t \delta_{m'm} \left(\delta_{l',l-1} \sqrt{l^2 - m^2} + \delta_{l',l+1} \sqrt{(l')^2 - m^2} \right) + \delta_{m'm} \delta_{l'l} (2l+1) \sigma_l \right] \frac{\eta_{l'm'}}{\sqrt{2l'+1}} = \delta_{m0} \frac{(2l+1)g^l}{\sqrt{4\pi}}.$$
(B20)

 μ_t

Suppose that the light in direction $\hat{\mathbf{s}} \in \mathbb{S}^2_{-}$ is detected at $\mathbf{r}_d \in \partial \Omega$. Here, \mathbb{S}^2_{-} denotes the set of unit vectors in outgoing directions. We have

$$I_{p}(\mathbf{r}_{d}, \hat{\mathbf{s}}) = \mu_{s} e^{i\mathbf{q}_{0} \cdot \rho} \sum_{l=0}^{l_{\max}} \sum_{m=-l}^{l} \eta_{lm} e^{-im\varphi_{\mathbf{q}_{0}}} Y_{lm}(\hat{\mathbf{s}}).$$
(B21)

4. GENERAL SOLUTION

Since the scattering phase function $p(\hat{\mathbf{s}}, \hat{\mathbf{s}}')$ depends only on $\hat{\mathbf{s}}$. $\hat{\mathbf{s}}'$, we can rewrite (3) as

$$p(\hat{\mathbf{s}}, \hat{\mathbf{s}}') = \sum_{l=0}^{l_{\max}} \sum_{m=-l}^{l} g^{l} Y_{lm}(\hat{\mathbf{s}}; \hat{\mathbf{k}}) Y_{lm}^{*}(\hat{\mathbf{s}}'; \hat{\mathbf{k}})$$
(B22)

for arbitrary $\hat{\mathbf{k}} = \hat{\mathbf{k}}(\nu, \mathbf{q})$. We note that

$$Y_{lM}^{*}(\hat{\mathbf{s}}; \, \hat{\mathbf{k}}) = \mathcal{R}_{\hat{\mathbf{k}}} Y_{lM}^{*}(\hat{\mathbf{s}})$$
$$= \sum_{m=-l}^{l} e^{im\varphi_{\hat{\mathbf{k}}}} d_{mM}^{l}(\vartheta_{\hat{\mathbf{k}}}) Y_{lm}^{*}(\hat{\mathbf{s}}).$$
(B23)

Let us express the eigenmodes as

$$\begin{split} \psi_{\nu}(\mathbf{r}, \hat{\mathbf{s}}, \mathbf{q}) \\ &= \sum_{l=0}^{l_{\max}} \frac{1}{\sqrt{\sigma_l}} \left\langle l | \phi_n(M) \right\rangle \, Y_{lM}(\hat{\mathbf{s}}; \hat{\mathbf{k}}) e^{-\mu_l \hat{\mathbf{k}} \cdot \mathbf{r}/\nu} \\ &= e^{i\mathbf{q} \cdot \boldsymbol{\rho}} e^{-\mu_l \hat{\mathbf{k}}_z(\nu q) z/\nu} \\ &\times \sum_{l=0}^{l_{\max}} \frac{\left\langle l | \phi_\nu \right\rangle}{\sqrt{\sigma_l}} \sum_{m=-l}^{l} e^{-im\varphi_{\hat{\mathbf{k}}}} d_{mM}^l(\vartheta_{\hat{\mathbf{k}}}) Y_{lm}(\hat{\mathbf{s}}). \end{split}$$
(B24)

We substitute the above $\psi_{\nu}(\mathbf{r}, \hat{\mathbf{s}}, \mathbf{q})$ in the homogeneous equation of the RTE. By using $l' = 0, ..., l_{max}$ and $m' = 0, \pm 1, \ldots, \pm l'$, we obtain

$$\sum_{l'=0}^{l_{\max}} \sum_{m'=-l'}^{l'} \frac{\langle l' | \phi_n(m') \rangle}{\sqrt{\sigma_{l'}}} Y_{l'm'}(\hat{\mathbf{s}}; \hat{\mathbf{k}}) \left(-\frac{\hat{\mathbf{s}} \cdot \hat{\mathbf{k}}}{\nu} + 1 \right) \mu_t$$
$$= \mu_s \sum_{l'=0}^{l_{\max}} \sum_{m'=-l'}^{l'} g_{l'}^{l'} \frac{\langle l' | \phi_n(m') \rangle}{\sqrt{\sigma_{l'}}} Y_{l'm'}(\hat{\mathbf{s}}; \hat{\mathbf{k}}).$$
(B25)

By rotating the reference frame in the inverse direction, we arrive at

$$\sum_{m'=-l_{\max}}^{l_{\max}} \sum_{l'=|m'|}^{l_{\max}} \frac{\langle l' | \phi_n(m') \rangle}{\sqrt{\sigma_{l'}}} Y_{l'm'}(\hat{\mathbf{s}}) \left(-\frac{\cos \vartheta}{\nu} + 1 \right) \mu_t$$
$$= \mu_s \sum_{m'=-l_{\max}}^{l_{\max}} \sum_{l'=|m'|}^{l_{\max}} g^{l'} \frac{\langle l' | \phi_n(m') \rangle}{\sqrt{\sigma_{l'}}} Y_{l'm'}(\hat{\mathbf{s}}).$$
(B26)

By multiplying $Y_{lm}^*(\hat{\mathbf{s}})$ $(-(l_{\max}-1) \le m \le l_{\max}-1, |m| \le l \le m$ l_{max}) on both sides and integrating over $\hat{\mathbf{s}} \in \mathbb{S}^2$, we obtain

$$\sum_{l'=|m|}^{l_{\max}} \left(b_{l+1}(m) \delta_{l+1,l'} + b_{l}(m) \delta_{l-1,l'} \right) \left\langle l' | \phi_n(m) \right\rangle$$
$$= \frac{\nu_n(m)}{\mu_n} \left\langle l | \phi_n(m) \right\rangle.$$
(B27)

In the above equation, we wrote $v = v_n(m)$. We see that $\frac{\nu}{\mu_t} = \nu_n(M)/\mu_t$ and $|\phi_n(M)\rangle$ are eigenvalues and eigenvectors of the following matrix-vector equation [8,9]:

$$B(M) |\phi_n(M)\rangle = \frac{\nu_n(M)}{\mu_t} |\phi_n(M)\rangle, \qquad (B28)$$

where $M = 0, \pm 1, ..., \pm (l_{\max} - 1)$, and matrix $B(M) \in \mathbb{R}^{(l_{\max} - |M| + 1) \times (l_{\max} - |M| + 1)}$ is a tridiagonal matrix whose elements are given by

$$\{B(M)\}_{ll'} = b_l(M)\delta_{l',l-1} + b_{l'}(M)\delta_{l',l+1}$$
(B29)

for $|M| \le l, l' \le l_{\text{max}}$. We used the notation such that $\langle l|B(M)|l+1\rangle = \langle l+1|B(M)|l\rangle = b_{l+1}(M)$. These $v_n(M)$ are approximate eigenvalues and discretized values of the continuous spectrum of Case's ν [34,37]. We note that for each pair of $v_n(M)$ and $\langle l | \phi_n(M) \rangle$ $(l = |M|, ..., l_{max})$, there exists a pair of eigenvalues $-\nu_n(M)$ and eigenvectors $(-1)^l \langle l | \phi_n(M) \rangle$ [8]. For the specific intensity $\psi(\mathbf{r}, \hat{\mathbf{s}})$ to vanish as $z \to \infty$, we take only $\lfloor (l_{max} - |M| + 1)/2 \rfloor$ eigenvalues and eigenvectors such that

$$\nu_n(M) > 0, \quad n = 1, 2, \dots, \left\lfloor \frac{l_{\max} - |M| + 1}{2} \right\rfloor.$$
 (B30)

From the point of view of the singular eigenfunction, the method of rotated reference frames is the spherical-harmonic expansion of the singular eigenfunction [10,37,38]:

$$\Phi_{\nu}^{m}(\hat{\mathbf{s}}) \approx \sum_{l=|m|}^{l_{\max}} \xi_{l}^{m}(\nu) Y_{lm}(\hat{\mathbf{s}}).$$
(B31)

Using $\langle \phi_n(M) | \phi_n(M) \rangle = 1$ and $\int_{\mathbb{S}^2} \mu | \Phi_{\nu}^m(\hat{\mathbf{s}}) |^2 d\hat{\mathbf{s}} =$ $2\pi \mathcal{N}^m(\nu)$ with the normalization factor $\mathcal{N}^m(\nu)$ from one-dimensional transport theory, we find

$$\xi_l^m(\nu) = \sqrt{\frac{2\pi\,\mu_t \mathcal{N}^m(\nu)}{\nu\sigma_l}} \,\langle l | \phi_n(m) \rangle \,. \tag{B32}$$

Furthermore, we note that $\langle l | \phi_{-\nu}(M) \rangle = (-1)^l \langle l | \phi_{\nu}(M) \rangle$ [8].

The specific intensity $\psi(\mathbf{r}, \hat{\mathbf{s}})$ is given by the superposition of eigenmodes $\psi_{\nu}(\mathbf{r}, \hat{\mathbf{s}}, \mathbf{q})$ with separation constant ν as

$$\psi(\mathbf{r}, \hat{\mathbf{s}}) = \frac{1}{(2\pi)^2} \sum_{\nu>0} \int_{\mathbb{R}^2} C_{\nu}(\mathbf{q}) \psi_{\nu}(\mathbf{r}, \hat{\mathbf{s}}, \mathbf{q}) d\mathbf{q}$$
$$= \frac{1}{(2\pi)^2} \sum_{\nu>0} \int_{\mathbb{R}^2} C_{\nu}(\mathbf{q}) e^{i\mathbf{q}\cdot\boldsymbol{\rho}} \sum_{l=0}^{l_{\max}} \sum_{m=-l}^{l} \frac{\langle l | \phi_{\nu} \rangle}{\sqrt{\sigma_l}} (-1)^m$$
$$\times e^{-im\varphi_{\mathbf{q}}} d_{mM}^{l} [i\tau(\nu q)] Y_{lm}(\hat{\mathbf{s}}) e^{-\mu_{\iota} \hat{k}_{z}(\nu q) z/\nu} d\mathbf{q},$$
(B33)

where $C_{\nu}(\mathbf{q})$ is determined later from the boundary condition. We note that

$$I_{\rm inc}^{(p)}(\mathbf{r}, \hat{\mathbf{s}}) = \mu_s e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}} \sum_{l=0}^{l_{\rm max}} \sum_{m=-l}^{l} \eta_{lm} e^{-im\varphi_{\mathbf{q}_0}} \times \left(R_{\mathfrak{n}}(\hat{\mathbf{s}} \cdot \hat{\mathbf{z}})(-1)^{l+m} - 1 \right) Y_{lm}(\hat{\mathbf{s}}).$$
(B34)

Let us find $C_{\nu}(\mathbf{q})$. As was done in [11], we introduce

$$\mathcal{B}_{ll'}^{m}(\mathfrak{n}) = \int_{S^{2}_{+}} R_{\mathfrak{n}}(\cos\vartheta) Y_{l'm}(\hat{\mathbf{s}}) Y_{lm}^{*}(\hat{\mathbf{s}}) d\hat{\mathbf{s}}$$

$$= \frac{1}{2} \sqrt{\frac{(2l+1)(2l'+1)(l-m)!(l'-m)!}{(l+m)!(l'+m)!}}$$

$$\times \int_{0}^{1} R_{\mathfrak{n}}(\mu) P_{l}^{m}(\mu) P_{l'}^{m}(\mu) d\mu.$$
(B35)

Note that $\mathcal{B}_{ll'}^{-m}(\mathfrak{n}) = \mathcal{B}_{ll'}^{m}(\mathfrak{n})$. Furthermore, we let $\mathcal{B}_{ll'}^{m}(\infty)$ denote $\mathcal{B}_{ll'}^{m}(\mathfrak{n})$, with $R_{\mathfrak{n}} = 1$. Let us take the Fourier transform for ρ and operate $\int_{\mathbb{S}^{2}_{+}} d\hat{\mathbf{s}} Y_{lm}^{*}(\hat{\mathbf{s}})$ on the boundary condition. By introducing $f_{Mn}(q)$ as

$$C_{\nu}(\mathbf{q}) = (2\pi)^2 f_{Mn}(q) \delta(\mathbf{q} - \mathbf{q}_0),$$
 (B36)

we obtain

$$\sum_{M=0}^{l_{\max}-1} \sum_{n=1}^{\lfloor (l_{\max}-|M|+1)/2 \rfloor} \left[\sum_{l'=\max(|m|,|M|)}^{l_{\max}} \left(\mathcal{B}_{ll'}^{m}(\infty) - (-1)^{l'+m} \mathcal{B}_{ll'}^{m}(\mathfrak{n}) \right) \frac{\langle l'|\phi_{n}(M) \rangle}{\sqrt{\sigma_{l'}}} \times \left(d_{mM}^{l'}(\vartheta_{\hat{\mathbf{k}}}) + (1-\vartheta_{M0})(-1)^{M} d_{m,-M}^{l'}(\vartheta_{\hat{\mathbf{k}}}) \right) \right] f_{Mn}(q)$$

$$= \mu_{s} \sum_{l'=m}^{l_{\max}} \eta_{l'm} \left((-1)^{l'} \mathcal{B}_{ll'}^{m}(\mathfrak{n}) - (-1)^{m} \mathcal{B}_{ll'}^{m}(\infty) \right)$$
(B37)

for $0 \le l \le l_{\max}$, $0 \le m \le l$. Note that equations for *m* and -m are the same, and hence, $f_{-M,n}(q) = (-1)^M f_{Mn}(q)$. Due to the fact that associated Legendre polynomials satisfy three-term recurrence relations, linearly independent equations are extracted from the above equations if equations with $l = m + 1 + 2\alpha \ (\alpha = 0, 1, \dots, \lfloor (l_{\max} - m - 1)/2 \rfloor)$ are taken for $m = 0, 1, \dots, l_{\max} - 1$.

Finally, we obtain

$$\psi(\mathbf{r}, \hat{\mathbf{s}}) = e^{i\mathbf{q}_0 \cdot \boldsymbol{\rho}} \sum_{l=0}^{l_{\max}} \sum_{m=-l}^{l} (-1)^m e^{-im\varphi_{\mathbf{q}_0}} Y_{lm}(\hat{\mathbf{s}}) K_{lm}(q_0, z),$$
(B38)

$$K_{lm}(q_0, z) = \sum_{M=-(l_{\max}-1)}^{l_{\max}-1} \sum_{n=1}^{\lfloor (l_{\max}-|M|+1)/2 \rfloor} f_{Mn}(q_0) \\ \times \frac{\langle l | \phi_n(M) \rangle}{\sqrt{\sigma_l}} d^l_{mM} [i\tau(v_n(M)q_0)] e^{-\mu_t \hat{k}_z z / v_n(M)}.$$
(B39)

Thus,

$$J_{+}(\mathbf{r}_{d}) = \int_{\mathbb{S}_{-}^{2}} (\cos \vartheta) I_{p}(\mathbf{r}_{d}, \, \hat{\mathbf{s}}) d\hat{\mathbf{s}} + \int_{\mathbb{S}_{-}^{2}} (\cos \vartheta) \psi(\mathbf{r}_{d}, \, \hat{\mathbf{s}}) d\hat{\mathbf{s}}$$
$$= -e^{iq_{0} \cdot \rho_{d}} A_{\text{RTE}}(q_{0}),$$
(B40)

where

$$A_{\text{RTE}}(q_0) = \sqrt{\pi} \sum_{l=0}^{l_{\text{max}}} (-1)^l \sqrt{2l+1} \left(\int_0^1 \mu P_l(\mu) d\mu \right) \times (\mu_s \eta_{l0} + K_l(q_0)) .$$
(B41)

Here,

$$K_{l}(q_{0}) = K_{l0}(q_{0}, 0) = \sum_{M \ge 0, n} f_{Mn}(q_{0}) \frac{\langle l | \phi_{n}(M) \rangle}{\sqrt{\sigma_{l}}}$$
$$\times (d_{0M}^{l}[i\tau(\nu_{n}(M)q_{0})]$$
$$+ (1 - \delta_{M0})(-1)^{M} d_{0, -M}^{l}[i\tau(\nu_{n}(M)q_{0})]).$$
(B42)

Note that $\int_0^1 \mu P_1(\mu) d\mu = \frac{1}{3}$, $\int_0^1 \mu P_l(\mu) d\mu = 0$ if l > 1 is odd, and when l is even,

$$\int_0^1 \mu P_l(\mu) d\mu = \frac{(-1)^{\frac{l}{2}+1} l!}{2^l (l-1)(l+2) [(\frac{l}{2})!]^2}.$$
 (B43)

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